Table 3

VLA and BIMA Continuum Results

Source		Freq.	Peak	rms	Spec. Index		RA (B1950)[1]	err[2]	DEC (B1950)[1]	err[2]	Ref.
		(GHz)	(mJy bm^{-1})	(mJy bm^{-1})	α	err	(hh mm ss.s)	(sec)	(° ′ ″)	″	
(1)	(2)	(3)	(4)	(5)	(6)	(7)	(8)	(9)	(10)	(11)	(12)
VLA Sources Near IRAS Objects											
RNO15FIR		8.46	$<0.12(3\sigma)$	0.04							
		4.86	$<0.24(3\sigma)$	0.08							
OrionA-W	A	22.215	2.8	0.5	1.15	0.19	05 30 14.415(2)	0.007	-05 37 50.82(4)	0.1	
		8.46	0.92	0.02							(a)
L1157		22.255	$<1.2(3\sigma)$	0.4							
	A	8.46	0.18	0.02	0.33	0.42	20 38 39.23(3)	0.05	67 51 35.8(2)	0.3	
		4.86	0.15	0.03							
B361		8.46	$<0.06(3\sigma)$	0.02							
		4.86	$<0.09(3\sigma)$	0.03							
L1251A		22.255	$<1.2(3\sigma)$	0.4							
	A	8.46	0.31	0.03	0.19	0.31	22 34 21.10(5)	0.1	75 01 34.1(1)	0.3	(b)
		4.86	0.28	0.04							
	B	8.46	0.19	0.03	>0.99		22 34 22.54(8)	0.1	75 01 37.6(2)	0.3	(c)
		4.86	$<0.11(3\sigma)$	0.04							
VLA Sources Outside IRAS Error Ellipse											
RNO15FIR	A	8.46	0.52	0.04	-2.26	0.16	03 24 26.052(7)	0.02	30 04 32.2(1)	0.3	
		4.86	1.81	0.08							
	B	8.46	4.46	0.04	-1.49	0.02	03 24 37.424(1)	0.02	30 01 23.71(1)	0.3	(d)
		4.86	10.17	0.08							
	C	8.46	2.78	0.04	-1.86	0.03	03 24 38.744(1)	0.02	30 01 11.02(2)	0.3	(d)
		4.86	7.79	0.08							
L1157	B	8.46	0.14	0.02	-0.98	0.35	20 38 14.11(4)	0.05	67 50 27.8(2)	0.3	
		4.86	0.24	0.03							
	C	8.46	0.78	0.02	-2.24	0.05	20 38 47.664(6)	0.05	67 49 05.34(4)	0.3	
		4.86	2.68	0.03							
L1251A	C	8.46	$<0.08(3\sigma)$	0.03	<-1.74						
		4.86	0.21	0.04			22 34 23.6	0.2	75 01 45.2	0.5	
	D	8.46	0.19	0.03	-0.18	0.45	22 34 26.88(8)	0.1	75 01 34.1(2)	0.3	
		4.86	0.21	0.04							
	E	8.46	1.39	0.03	1.02	0.03	22 34 18.78(1)	0.1	75 04 10.13(3)	0.3	
		4.86	0.80	0.04							
BIMA Sources Near IRAS Objects											
L1157	C1	107.1	36.3[3]	3.9			20 38 39.34	0.2	67 51 34.6	0.8	(e)
L1251A		107.1	$<11.7(3\sigma)$	3.9							

Notes - (1) The astrometric resolution of the last digit follows in parenthesis.
(2) The absolute error calculated from the prescription in Sec. 2.2.
(3) Integrated emission at a level of 64 ± 18 mJy is observed from an area of 7.″3 x 5.″9, PA=105°.
References - (a) See also Tofani *et al.* 1995; (b) RD-D, Rosvick and Davidge 1995; (c) RD-A, Rosvick and Davidge 1995; (d) see also Schwartz *et al.* 1985; (e) see also Gueth *et al.* 1997.